

**Ch 1 Outline of Universe**

Hierarchy: nuclei/atoms/stars/galaxies  
 Solar System  
 Nearby Stars, Our Galaxy  
 Galaxies and the visible Universe  
 Lengths: AU, light year, parsec  
 Angles: degree, arcmin, arcsec

**App.A** Units, Powers of 10

**Ch 6.1 Nature of Light**

Speed of light =  $c = 300,000$  km/s.  
 Spectrum (blue = 400nm, red = 600nm)  
 Electromagnetic waves  
 Electromagnetic Spectrum:  
 Radio, microwave, infra-red, optical,  
 ultra-violet, X-ray, gamma-ray  
 Atmospheric windows (transparency/opacity)  
 Wavelength ( $\lambda$ ), frequency ( $f$ ),  $c = f\lambda$   
 $f$  measured in Hertz (Hz) ie "per sec"  
 Photons:  $E = hf = hc/\lambda$   
 $h =$  Planck's constant =  $6.6 \times 10^{-34}$  J s

**Ch 7 Creation of light:** accelerated charge

Thermal radiation: hot solids, liquids, dense gases  
 Temperature scales: Kelvin (K),  
 Celsius (C), Fahrenheit (F) [0 K = -273 C]  
 Blackbody (perfect) emitter/absorber  
 Blackbody spectrum shape  
 Stars (dense gas) approximate blackbody spectra  
 Higher Temp, shorter (bluer) peak,  
 $\lambda_{\text{peak}}(\text{nm}) = 3 \times 10^6/T$  (Wein's Law)  
 Star color  $\rightarrow$  star temperature  
 Higher temp, more emission  
 $E/\text{unit area} = \sigma T^4$  (Stefan-Boltzmann Law)  
 Line radiation created by atoms  
 Atomic structure: protons, neutrons, electrons  
 protons and neutrons in nucleus  
 electrons in orbits  
 # of protons  $\rightarrow$  element  
 # of neutrons + protons  $\rightarrow$  isotope  
 # electrons = # protons (neutral atoms)  
 # electrons  $\neq$  # protons (ions)  
 Bohr picture of electron orbits  
 only certain permitted orbits  
 orbit levels  $\rightarrow$  energy levels  
 electrons can jump between levels  
 jump up  $\rightarrow$  absorption of photon  
 jump down  $\rightarrow$  emission of photon  
 jump out (lost)  $\rightarrow$  ionization  
 Hydrogen atom energy levels ( $n=1,2,3\dots$ )  
 Lyman ( $n=1$ ), Balmer ( $n=2$ ), etc series of lines  
 Kirchoff's Laws: conditions for  
 creating Emission/Absorption lines  
 Stellar spectra  
 Absorption lines from photosphere  
 Hydrogen Balmer series changes in strength  
 weak at low temp: all in ground state ( $n=1$ )

weak at high temp: all ionized  
 strongest at  $\sim 10000$  K

Classification of absorption spectra

O B A F G K M

subdivided by #s 0 to 9

Temperature (spectral) sequence

O stars 50,000 °K, M stars 2000 °K

T decreasing: ions  $\rightarrow$  atoms  $\rightarrow$  molecules

Composition of stars: 74% H, 24% He, 2% all others  
 abundances decrease for heavier elements

Doppler shift:  $v/c = (\lambda - \lambda_o)/\lambda_o$

$v$  is radial velocity (RV) only

redshift (away), blueshift (towards)

no shift (transverse)

Line broadening:

high pressure  $\rightarrow$  wide lines.

**Ch 8 The Sun:** an average star

Global properties

size, mass, density, luminosity,

temperature, composition, rotation

Photosphere ('surface')

Granulation (convection cells)

Chromosphere (He discovery)

Corona (hot)

Solar wind  $\rightarrow$  aurora, comet tails

Sunspots, sunspot cycle

Magnetic field, magnetic cycle

Prominences, flares

Solar interior: helioseismology

The Sun's Energy Source & Structure

Nuclear binding energy curve (fusion/fission)

Nuclear (hydrogen) fusion, p-p chain

$E=mc^2$ : energy from mass

$4H \rightarrow {}^4\text{He} + \text{Energy}$  (0.7% efficient)

Solar neutrinos; detectors; discrepancy

**Ch 9 Stars**

Star distances:

stellar parallax ( $p$ )

$d(\text{pc}) = 1/p(\text{arcsec})$

1pc = 3.26 light years

difficult beyond 1000 pc

Star motion (few arcsec/century; few km/s):

proper (transverse) & doppler (radial)

Star luminosity,  $L$ , intrinsic (Watt)

Star brightness,  $b$ , at earth (Watt/m<sup>2</sup>)

brightness depends on distance and luminosity

inverse square law:  $b \propto \frac{1}{d^2}$ ;  $b = \frac{L}{4\pi d^2}$

e.g. Sun:  $L = 3.8 \times 10^{26}$  Watt,  $b = 1400$  Watt/m<sup>2</sup>

Sirius:  $L = 9.1 \times 10^{27}$  Watt,  $b = 1.14 \times 10^{-7}$  Watt/m<sup>2</sup>

Apparent magnitude,  $m$

Hipparchus, 1  $\rightarrow$  6 (brighter  $\rightarrow$  fainter)

difference of 5 mag is  $\times 100$  in brightness

difference of 1 mag is  $\times 2.512$  in brightness

absolute magnitude ( $M$ ) =  $m$  at 10pc

Star sizes: large range

surface area of star (sphere) =  $4\pi R^2$

Luminosity of star:  $L = 4\pi R^2 \sigma T^4$

H–R diagram (Luminosity vs Temp)

Main sequence (MS), giants, white dwarfs

Luminosity/size class (width of absorption lines)

large R → low pressure → narrow lines

→ Luminosity Class (LC) & star size

LC I, III, V = supergiants, giants, main sequence

e.g. Sun: G2V, Arcturus: K2III

Binary stars

visual; spectroscopic; eclipsing

Star Mass determination from binary orbits:

stars orbit about center of mass

$k (M_1 + M_2) = \frac{a^3}{P^2}$  (Kepler's 3<sup>rd</sup> Law)

$k = 1$  if M in  $M_\odot$ ,  $a$  in AU,  $P$  in years

mass ratio  $M_1/M_2 = V_2/V_1$

Mass–Luminosity relation for MS stars

$L \propto M^{3.5}$ , eg M0 =  $0.5 M_\odot$ ; A0 =  $4 M_\odot$

Mass–Lifetime relation for MS stars ( $t \propto M^{-2.5}$ )

Mass limits: max  $\sim 50 M_\odot$ , min  $\sim 0.08 M_\odot$

Luminosity function (census by brightness):

lower luminosity stars more common

night sky stars usually luminous

### Ch 10 Interstellar medium (ISM): low density gas & dust

Galactic “atmosphere”,  $\sim 5\%$  mass in stars

Dust: composition/absorption/reddening

cold, so thermal emission in IR

Three kinds of nebulae: dark/reflection/emission

HI: electron spin flip → 21 cm (radio) photon

Molecules: e.g. CO “visible” in radio (not H<sub>2</sub>)

X-ray emitting (hot) gas from supernova explosions

four phases (approximate pressure balance):

1) cold ( $\sim 20\text{K}$ ) “dense” ( $\sim 10^3/\text{cm}^3$ ) molecular clouds

2) warm ( $\sim 500\text{K}$ ) neutral hydrogen, HI clouds

3) hot ( $\sim 10^4\text{K}$ ) ionized hydrogen, HII regions

4) v. hot ( $\sim 10^6\text{K}$ ), low density ( $0.01/\text{cm}^3$ ), pervasive

cycle: gas → stars → gas

### Ch 11 Star Birth

Gas (molecular) clouds usually supported:

thermal+magnetic pressure/turbulence/rotation

Can collapse if triggered by shock

supernovae; stellar winds; spiral arms; cloud collision

core heating from gravity: “protostars”

hidden cocoon phase (IR sources), disks & jets

planets form; discovering extrasolar planets

Hydrogen fusion begins: star is “born”

evolutionary tracks on H-R diagram

Zero Age Main Sequence (ZAMS)

location on MS mass dependent (O high, M low)

formation time mass dependent (O short, M long)

Brown dwarf: failed star,  $M_* < 0.08 M_\odot$ , no fusion

HII regions, O&B stars ionize H, pink (H $\alpha$ ) color

Star winds & supernovae remove gas → young star cluster

Star Structure:

Hydrostatic equilibrium:

gravity inwards = pressure outwards

Energy source: p–p chain; CNO cycle

Energy transport:

radiation, convection, (conduction)

regions depend on mass

Star stability: stellar thermostat

### Ch 12 Star Evolution

Computer model of Sun’s interior:

centrally concentrated (core/envelope)

nuclear reactions in core ( $< 20\%$ )

inner 80% radiative

outer 20% convective

Life on the Main Sequence (MS)

core H fusion ( $4\text{H} \rightarrow \text{He}$ )

p–p chain (low mass), CNO cycle (high mass)

longest stage of life (eg sun  $10^{10}$  yrs)

MS lifetime: Lower mass → Longer lifetime

Leaving MS, up Red Giant Branch (RGB) 1<sup>st</sup> time

H fuel exhausted in core

core contracts & heats

H burns in shell, powerful

envelope expands & cools

Red Giant ( $R_* \sim 1\text{AU}$ ), luminous ( $L_* \sim 10^3 L_\odot$ )

Beyond RGB

He ignites in core,  $3\text{He} \rightarrow \text{C}$  (triple alpha)

( $M_* < 2 M_\odot$  Helium flash)

lands on “Horizontal Branch” (HB)

He fuel exhausted in core

He & H shells burning, very powerful

envelope expands and cools

up RGB 2<sup>nd</sup> time (AGB)

red supergiant ( $R_* \sim 5\text{AU}$ ,  $L_* \sim 10^4 L_\odot$ )

Star Clusters: groups of the same age

Globular clusters: old, halo, metal poor

Open clusters: young, disk, metal rich

Cluster H–R diagrams: isochrones (same age)

Young cluster isochrones

Main Sequence Turn Off point (MSTO)

measuring cluster age from MSTO

Variable stars

instability strip on H–R diagram

stellar pulsation, periods 1 – 90 days

RR Lyrae (fainter, lower mass)

Cepheids (brighter, higher mass)

period–luminosity relation

distance estimates

### Ch 13 Star Death

Low mass stars ( $< 0.4 M_\odot$ ):

long lives; cannot ignite He

Intermediate mass stars (eg the Sun):

Large size → low surface gravity → mass loss

expanding envelope = “Planetary Nebula” (PN)

reveals hot core → White Dwarf (WD) star

UV ionizes PN → emission line glow

WD *slowly* cools → black dwarf

High mass stars ( $M_* > 10M_\odot$ ) DIFFERENT

- short MS lifetime
- heavier elements burn (He→C/O→Ne... etc)
- higher core temps & densities, shorter durations
- onion shell structure, core + shells
- iron core: no more energy (minimum in mass defect)
- core collapse (0.2 seconds; radius 3000km → 10km)
- protons + electrons → neutrons + neutrinos
- huge neutrino release
- core bounce → rising shock → star explodes

Type II supernova (eg SN 1987A)

- total energy  $\sim 10^{46}$  Joules (gravity origin)
- energy emerges: 99% neutrinos, 1% kinetic, 0.01% light
- neutrino burst detected from SN 1987A
- light outshines galaxy for a week
- fades with decay  $^{56}\text{Co} \rightarrow ^{56}\text{Fe}$  ( $t_{\frac{1}{2}} = 77$  days)
- expanding shell (supernova remnant)

Nucleosynthesis (origin of elements)

- $\alpha$  elements; r & s process (neutron addition)
- recycle into ISM → new stars (and planets and us!)
- gradual element enrichment of galaxy

#### Ch 14 Star Corpses

White dwarf (WD)

- dead core of intermediate mass star
- earth size, carbon/oxygen, hot (white), faint
- density  $\sim 5 \times 10^5$  g/cm<sup>3</sup>
- electron degenerate pressure support
- degeneracy pressure independent of temperature
- more massive WDs are smaller
- limiting mass  $1.4M_\odot$  (Chandrasekhar mass)

Mass accretion onto WD from companion:

- nova: hydrogen detonation of surface layer
- Type I supernova: carbon detonation of entire star

Neutron stars:

- remaining core after Type II Supernova
- size  $\sim 10$ km, density  $\sim 10^{14}$  g/cm<sup>3</sup>
- neutron degenerate pressure support
- mass limit about  $3M_\odot$

Pulsars:

- rotating magnetized neutron stars
- conservation of angular momentum → rapid spin
- lighthouse model
- Crab nebula powered by pulsar (synchrotron)
- slowing down (young pulsars spin faster than old)

Mass accretion onto NS from companion:

- He detonation of surface → X-ray burster
- rapid spin-up → millisecond pulsars

Black Holes:

- escape velocity,  $V_{esc} = \sqrt{2GM/R}$
- for  $V_{esc} = c$ , we get  $R_s = 2GM/c^2$
- $R_s$  = Schwarzschild radius = event horizon =  $3M/M_\odot$  km
- $M > 3M_\odot$  collapse unstoppable → singularity

Near BH: time slows, light redshifted, strong tides

Spinning (Kerr) black holes, drag space around.

Searching for black holes:

rapid X-ray variability → small size

Cyg X-1, accreting binary,  $M_{BH} \sim$  few  $M_\odot$

#### Ch 15 Our Galaxy

Appearance of Milky Way (MW) → slab geometry

Star counts → sun at center (Herschel)

Globular cluster distribution → sun off-center (Shapley)

Discrepancy: dust absorption in ISM (Trumpler)

Galaxy shape:

- radius  $\sim 15$ kpc, sun  $\sim 8.5$ kpc from center
- disk of stars + ISM, bulge, nucleus,
- large spherical halo with globular clusters

Motions:

- Halo orbits: elliptical; random orientation
- Disk orbits (eg sun):  $\sim$ circular in disk plane
- sun's speed  $\sim 200$  km/s
- sun's orbital period  $\sim 200$  million yrs
- about 70 rotations since birth of galaxy
- Kepler's law →  $M \sim 10^{11}M_\odot$  inside sun

Rotation curve:

- rapid rise, then constant (at  $\sim 200$  km/s)
- differential rotation (inner stars overtake)
- dark matter extends beyond stars (unknown)

Stellar Populations:

- Pop I: young, disk, enriched by heavy elements
- Pop II: old, halo, depleted in heavy elements

Galaxy Formation:

- collapse of protogalactic cloud
- + interactions/accretion of dwarf companions

Spiral structure:

- seen in HI 21cm maps
- seen in distribution of star formation regions
- arms follow star formation → brighter/bluer

Theories of spiral origin

- NOT simple rotation → windup too tight
- self-propagating star formation
- density wave theory: slow pattern speed
- density wave triggered by bar or passing neighbor

Galactic Nucleus, dense star cluster, star formation

- star orbits traced, periods  $\sim 10$  yrs
- black hole with mass  $\sim 2.6 \times 10^6 M_\odot$

#### Ch 16 Other Galaxies

Early work (1850s): Lord Rosse → Nebulae

Shapley–Curtis debate: nature of spiral nebulae

Cepheids in M31 → external galaxy (Hubble)

Distribution & Total Number:

- Slod Survey of  $\sim 10^6$  galaxies: walls/sheets/voids
- Hubble Deep Field →  $10^{11}$  galaxies total

Galaxy shapes: Hubble's classification

- tuning fork diagram
- spirals: Sa→c: big→small bulge, tight→open arms
- barred spirals: SBa,b,c ; have inner bar of stars
- ellipticals: smooth, round (E0)→ flat (E7), no ISM
- lenticular: like E plus disk but no arms or ISM
- irregulars (often smaller)

Galaxy distances

Cepheid variables (ok to  $\sim 1\text{Mpc}$ )  
 further, need bright “standard candles”  
 eg brightest stars, HII regions, supernovae  
 Velocity of galaxies (almost all redshifts)  
 Hubble diagram: velocity (km/s) *vs* distance (Mpc)  
 straight line,  $v \propto d$ , or  $v = Hd$  (Hubble Law)  
 H: Hubble’s constant  $70 \pm \text{km/s/Mpc}$   
 (nearby galaxies don’t follow the law)  
 use to estimate distance ( $v \rightarrow d$ )  
 caused by universe expanding  
 Galaxy Properties: large range in size, mass, luminosity  
 Galaxy masses and mass/light ratios  
 rotation curves  
 dark (missing) mass  
 Galaxy interactions/collisions/mergers:  
 gravitational perturbations  $\rightarrow$  distortion  
 gas clouds collide  $\rightarrow$  strong star formation  
 no stars collide (too much space)  
 spiral + spiral  $\rightarrow$  elliptical  
 Groups of galaxies: eg Local Group (3 spirals; 20 dwarfs)  
 Clusters of galaxies:  
 can contain thousands of galaxies  
 few big galaxies, many more smaller ones  
 often large central galaxy  $\rightarrow$  cannibalized others  
 Doppler shifts  $\rightarrow \sim 500 \text{ km/s} \rightarrow 10^8 \text{ yr}$  to cross  
 velocities  $\rightarrow$  mass  $\rightarrow$  dark mass dominates  
 gravitational lenses  $\rightarrow$  dark mass dominates  
 Hot ( $10^7\text{K}$ ) hydrostatic atmosphere  $\rightarrow$  dark mass dominates  
 Galaxy Evolution  
 study distant (therefore younger) galaxies  
 find smaller building blocks (resemble LMC)  
 mergers  $\rightarrow$  star formation, chemical enrichment  
 Hierarchical assembly: smaller  $\rightarrow$  bigger

**Ch 17 Active Galaxies**

Seyfert galaxies: bright variable nucleus  
 Radio galaxies: jets and lobes  
 Synchrotron emission: fast electrons in magnetic field  
 Supermassive black holes ( $10^7 - 10^9 M_\odot$ ) in galactic nuclei  
 Accretion disk releases gravitational energy  $\rightarrow$  luminous  
 High speed bipolar jets often created  
 jets can appear superluminal  
 Even non-active galaxies may have nuclear black holes  
 Activity triggered by galaxy interactions  $\rightarrow$  gas inflow  
 Quasars:  
 Discovery history:  
 bright radio sources but faint galaxies or “stars”  
 (“quasars”: quasi-stellar radio source)  
 redshifts  $\rightarrow$  very distant  $\rightarrow$  VERY luminous  
 some brighter than 100 galaxies  
 variable ( $\Delta t \sim \text{months}$ )  $\rightarrow$  tiny source ( $R < c\Delta t$ )  
 energy source puzzle  $\rightarrow$  distances doubted  
 Evolution; quasars more common in past (further away)

**Ch 18 Cosmology: The Big Bang**

Global Properties: isotropy; homogeneity;  
 expansion; universality of physical law

Hubble law: cosmic expansion (cake analogy)  
 linear expansion, gradient  $\rightarrow$  H (Hubble’s constant)  
 Cosmological (not Doppler) redshift: spacetime expands  
 No center: everyone sees same expansion  
 All together at one time  $\rightarrow$  Big Bang  
 Time since BB:  $t \sim d/v \sim 1/H$  eg  $H=70 \rightarrow t \sim 13 \text{ Gyr}$   
 Lookback time = light travel time (eg  $1\text{Gly} \rightarrow 1\text{Gyr}$  ago)  
 Current limits are 80% lookback to BB (adolescent)  
 Particle horizon at 100% lookback to BB  
 Shape & size of universe:  
 matter curves spacetime (curvature  $k = +1, 0, -1$ )  
 curvature depends on average density  
 critical density ( $k=0$ )  $\sim 9.5 \times 10^{-27} \text{ kg/m}^3$   
 measure  $k$ , use cosmic triangle  $\rightarrow k=0$  (“flat” geometry)  
 Contents:  
 Atoms (4%) Dark Matter (23%) Dark Energy (70%)  
 total adds to critical density ( $k=0$ )  
 Microwave background:  
 whole sky: thermal spectrum,  $T=2.7\text{K}$   
 some distortion: earth’s motion; MW foreground  
 Penzias & Wilson (1963) ..... WMAP (2003)  
 recombination in  $3000\text{K}$  gas; hot fog cools & clears  
 redshift factor  $\sim 1000$ ;  $380,000 \text{ yrs}$  after BB  
 First million years & Sound:  
 hot thin glowing gas  $\rightarrow$  sound waves present  
 caused by gravity pulling on gas  $\rightarrow$  dropping pitch  
 Accelerating Expansion:  
 measure expansion history using distant SN  
 slower in past  $\rightarrow$  expansion speeding up  
 explain using Einstein’s  $\Lambda$  parameter (vacuum energy)  
 gets more important  $\rightarrow$  Universe expands for ever  
 Early History  
 $t \sim 50,000\text{yr}$ : matter density = radiation density  
 $t \sim \text{few minutes}$ : conditions = stellar interior  
 nucleosynthesis: H fusion to He  
 predict 24%; observe 24%  $\rightarrow$  evidence for hot BB  
 Deuterium abundance  $\rightarrow$  baryon density ( $\sim 4\%$ )  
 $t < 1 \text{ second}$ :  
 high Temp  $\rightarrow$  matter created from particle collisions  
 threshold temps: electrons ( $10^{10}\text{K}$ ) freeze out at  $1\text{sec}$   
 protons ( $10^{12}\text{K}$ ) freeze out at  $100 \mu\text{s}$   
 protons break into quarks at  $T > 10^{13}\text{K}$  ( $t < 1\mu\text{s}$ )  
 Four eras: Quark; Hadron; Lepton; Radiation  
 Origin of Forces  
 four forces: gravity, weak, EM, strong.  
 strength/identity is temperature dependent  
 EM & weak join at  $10^{-10}\text{s}$  ( $10^{15}\text{K}$ );  
 add strong at  $10^{-35}\text{s}$ ; add gravity at  $10^{-43}\text{s}$   
 Problems with standard BB theory:  
 Problems: Flatness; Horizon; Structure; Matter.  
 Inflation may solve all these  
 huge early expansion  $\rightarrow k=0$ ; smooth on large scales  
 quantum roughness  $\rightarrow$  structures  
 GUT transition favour matter over antimatter  
 Exciting times !